IRSIN INSTITUT DE RADIOPROTECTION ET DE SÛRETÉ NUCLÉAIRE

Faire avancer la sûreté nucléaire

Preliminary studies for ASTRID-like SFR implementation in the CLASS code

2nd Technical Workshop on Fuel Cycle Simulation

July 19th - 21st 2017

Columbia, South Carolina U.S.A.



Neutronics and Criticality Safety Department

19/07/2017

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### Outline

### **Objectives**

Simulation tools

**Geometry modeling** 

**Preliminary results** 

- Static calculations
- Depletion calculations

#### **Conclusion & Perspective**



### Objectives

#### Long term objectives

- Scenario simulations (integrating ASTRID-like SFR)  $\rightarrow$  CLASS code
- Estimation of the accuracy associated with Physic models
  - Fuel production & evolution
- Databank (K<sub>eff</sub>, Flux, XS): many depletion calculations

#### Medium term objectives

- Many full core calculations: static & depletion simulations
- ASTRID-like SFR: 2 types of axially heterogeneous assembly
  - Huge cost in term of CPU time & memory resources

#### Short term objectives

- Assembly calculations: full core preparation → tools & methods
- Accuracy & CPU time optimization
  - Methods used for full core calculations



Core Library for Advanced Scenario Simulation



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Core Library for Advanced Scenario Simulation



### Simulation tools

Monte-Carlo calculations: MORET 5.D.1 Beta

- Criticality code
- Soon available at OECD

#### Depletion calculations: VESTA 2.2 Beta

- Thin multigroup calculation (43 000 groups)
- Version 2.1.5 available at OECD
- Nuclear data: JEFF.3.1

#### Parametrical calculations launching: Prométhée

- Parallel distribution of calculations
- Algorithms for advanced engineering based on R language
- Post-processing: R scripts
  - (Library plotly, rCharts, devtools)









### Geometry modeling

#### **7** With the MORET code

#### 2 types of hexagonal assembly

- External: 1 fissile area, 1 fertile area
- Internal: 2 fissile areas, 2 fertile areas
- Depletion: 1 fuel / zone / assembly
- 217 identical fuel pins per assembly
- Helicoidally spacer wires: not modelled
- Semi-infinite calculation





### Geometry modeling

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Sodium

-0.214

**Y** 0.428

0.214

0.0

Х

Semi-infinite calculation



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0.214

0.428

> Void

### Preliminary results



### Static calculations: internal assembly

### **7** Experimental plan

#### Simulation parameters

- 10 inactive cycles (choice)
- Active cycles
- $\rightarrow$  [20 , 50 , 100 , 150 , 200 , 500]
- Number of source particles per zone
- $\rightarrow$  [1 , 5 , 10 , 15 , 20 , 30 , 50 , 100]

#### Number of particles

 $N_p \times N_{ca} \times N_{pins} \times N_{zones}$ 

#### A priori selected criteria

- CPU time < 10 min</p>
- Uncertainty on  $K_{eff} \sim 1.5\%$  Values on full
- Uncertainty on Flux ~ 1%
   assembly

#### Same parameters for both assembly

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#### Parallel plot of the simulations





### Static calculations: internal assembly



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- Uncertainty on Flux ~ 1%

Values on ful assembly

### Static calculations: internal assembly

#### Parallel plot of the simulations n CA n tot CPU.time std Keff std Flux n P Keff Flux 100 500 0.0025 -3.07 40,000,000 -198.5 1.066 2.5-80 400 -0.0020 600· 198.0 30,000,000 -1.064 -2.0-60 0.0015 -300 -197.5 400 1.5 -20,000,000 -1.062 197.0 40 0.0010 200 -1.0-200 196.5 -10,000,000 1.060 20 0.0005 100 -0.5 -196.0

A priori selected criteria

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assembly

Selected parameters: 50 active cycles Internal: **External:** 4340 part/cycle 2170 part/cycle 9.03 min 2.35 min

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### Depletion calculations: internal assembly

#### **7** Experimental plan

#### Simulation data

- Burnup: 110 GWd/t
- Mean power: 5.12 MWth
- Irradiation time: 2817 d

#### Modelling parameters

- Number of timesteps
- $\rightarrow$  [1, 4, 6, 8, 10, 15, 25, 50, 75, 100]
- Propagation of statistical Monte-Carlo uncertainty (random seed)
- → 50 seeds [1000: 246000]

#### Quantity of interest

- Reaction: Fission, Capture, (n,2n)
- Isotope: <sup>238</sup>U, <sup>239</sup>Pu, <sup>237</sup>Np, <sup>243</sup>Cm, ...
- K<sub>eff</sub> & associated uncertainty, Flux & Power, XS, Composition



### **Depletion calculations:** internal assembly

0.99

0 00 0.992

0.99 0.988

#### Experimental plan

#### Simulation data

- Burnup: 110 GWd/t
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#### ₩ 0.986 0.984 0.982 d<sup>.98</sup> 0.978

0.976 0.974 0.972 95 100 105 110 115 120 125 Burnup (MWd/kgHM)

#### 100 timesteps: reference calculation to estimate differences due to the number of timesteps

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### Reference calculation for timesteps impact

Evolution of Keff(BU) pour 100 & 200 timesteps

System\_100S5

•••• System\_200S5

# Example of K<sub>eff</sub> Impact of the timestep number



Deviation between the reference polynomial interpolation and for other timesteps







Envelope of the evolution of Keff(BU) for 10 timesteps & different random seeds



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### **オ** Example of K<sub>eff</sub>



10 timesteps:

0.29% of relative difference compared to 100 timesteps

0.38% of relative error due to random seed

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#### Impact of the timestep number

Maximal polynomial difference compared to the reference



#### A priori selected criteria

- Uncertainty on  $K_{eff}$  ~ 1.5‰
- Uncertainty on Flux ~ 1%

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#### Impact of the timestep number

Maximal polynomial difference compared to the reference



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#### Deviation due to timestep number variation



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- Uncertainty on  $K_{eff}~$  ~ 1.5‰
- Uncertainty on Flux ~ 1%

# From the terms of the seed Seed

#### Relative error for 10 timesteps

- 0.38% on the  $\rm K_{\rm eff}$
- 2.25% on Flux, Upper Fissile zone
- 1.83% on Flux, Internal Fertile zone
- 2.94% on Flux, Lower Fissile zone
- 4.26% on Flux, Fertile Blanket



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*Error*<sub>random seed</sub> > *Difference*<sub>timesteps</sub>



### Depletion calculations: XS & Composition

#### Same method for other quantities of interest

- Similar trends: no visible gain for  $timesteps \ge 10$
- Error<sub>random seed</sub> generally superior than Difference<sub>timesteps</sub>

#### Additional observation

#### Relative error due to the random seed for 10 timesteps



CompositionIsotopeUpper Fissile zone EOCFertile Blanket EOC238U0.11 %0.08 %239Pu0.13 %0.77 %237Np2.52 %5.59 %243Cm0.75 %4.83 %

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### **Conclusion & perspective**

#### Final objective

- Integration of ASTRID-like SFR into CLASS code
- Utilisation of physic models based on databanks
- Accuracy of model predictions  $\rightarrow$  Errors associated with the quantities of interest

#### Objective of this preliminary study

 Development of tools & methods to estimate errors at each step of the databank generation

#### Conclusion & perspective of this preliminary study

- Might be sufficient to use a relatively low number of timesteps
- Increase of the statistical error with 10 timesteps during the depletion calculation
- Under-estimation of the number of neutrons during the static calculation
- Trade-off between CPU time and number of neutron increase
- $\rightarrow$  Iteration of the method with more particle to confirm previous conclusion
- $\rightarrow$  Estimation of error due to the random seed for 100 timesteps



### **Conclusion & perspective**

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### Thank you for your attention

